s/133/62/000/006/012/015 A054/A127

Continuous electrolytic pickling of ...

plied instead of lead ones. The pickling of a 1.6-mm diameter 1Kh18N9T steel wire at current densities of 34 - 42 a/dm2 for 12 seconds yielded the optimum result, a bright, clean surface. When applying the possible maximum current density for the given conditions (42 a/dm2) a satisfactory surface was obtained in 9.6 seconds. Electrolyte C) gave results similar to A); electrolyte B) was unsatisfactory. Increasing the current density above 40 a/dm² did not accelerate the process: the required time could not be shortened under 12 seconds. The alkaline solution and electrolytes A) and C) can also be applied in pickling carbon steel wire. In that case, at a current density of 15 - 18 a/dm2 the output of the process increases by a factor of 1.6 - 1.7 as compared with the continuous chemical process. The industrial-scale tests were carried out by setting 90-mm wide baths of 1Kh18N9T steel in the conventional thermal pickling equipment, filled with the at 460 - 470°C following solutions:

65% NaOH + 30% NaNO3 + 5% NaCl at 80 - 85°C 18% H₂SO₄ + 1% NaCl + 5% NaNO₃ Alkaline bath:

. at room temperature Acidic bath:

In pickling 3.6 mm diameter OKh18N9T and Kh18N11M steel wires, a clean, bright and scale-free surface was obtained at a rate of 10.5 m/min, (20 sec. in the al-

Card 2/3

Continuous electrolytic pickling of ...

S/133/62/000/006/012/015 A054/A127

kaline, 29 sec. in the acidic and 11 sec. in the bleaching bath). In the simultaneous pickling of 24 wires, at a rate of 20 ton/day, a generator power of 140 kw (51 v, 2,750 a) is required. In pickling stainless steel strips, (0.5 mm thick, 40 mm wide) in the pilot installation, the cathode and anode plates (0.6 mm long in the alkaline bath and 1 m long in the acidic bath) were set parallel to the movement of the strip. On account of the larger surface of the strip, the maximum current density was lowered to 12 - 15 a/dm². The strips tested were made of 1kh18N9T, 1kh18N9, 34 432 (E1432), 34 435 (E1435), X13H4F9 (kh13N4G9) and 1X13 (1kh13) steel grades. In the alkaline bath 100-% NaOH (at 450°C), in the acidic bath solutions A) and C) (at 70°C) were tested. A satisfactory surface was obtained with these solutions, when keeping the strip 6 - 9 seconds in the alkaline bath and 9 - 15 seconds in the acidic one. As in currentless continuous chemical pickling - under industrial conditions - the pickling of the same strip requires 82 seconds, the electrolytic method increases the output of the process 3 - 4 times. For pickling strips 0.5 mm thick and 400 mm wide at a current density of 15 a/dm² and with electrode plates 7 m long, the generator power required will be 1,260 kw (150 v, 8,400 a). There is 1 figure.

ASSOCIATION: Zavod "Serp 1 molot" ("Serp 1 molot" Plant)

Card 3/3

ZHETVIN, N.P., kand.tekhn.nauk; ZUEREV, I.P., inzh.; BOLOUSOV, A.S., inzh.;
ANDRATSKIY, K.K., inzh.; DIQMIDOVA, A.A., inzh.

Treatment of rope wire in continuous heat treating and pickling units. Stal' 22 no.7:661-663 Jl '62.

(Wire industry)

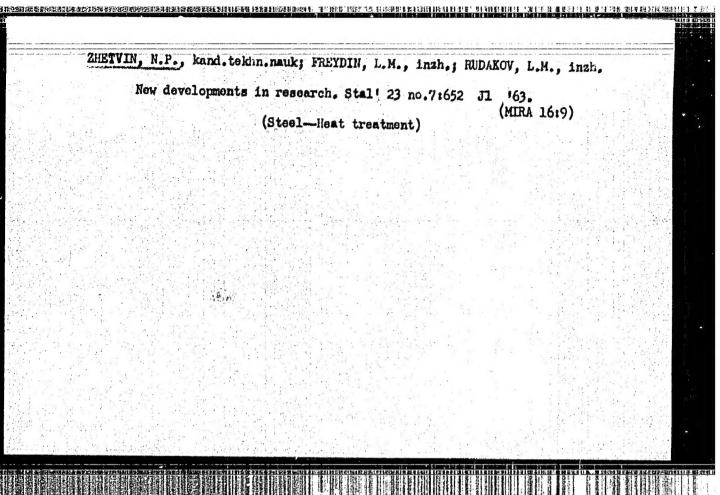
ZHETVIN, N.P., kard.tekhn.nauk; GORBATENKO, I.V., inzh.; KONTSEVAYA, Ye.H., irsh.

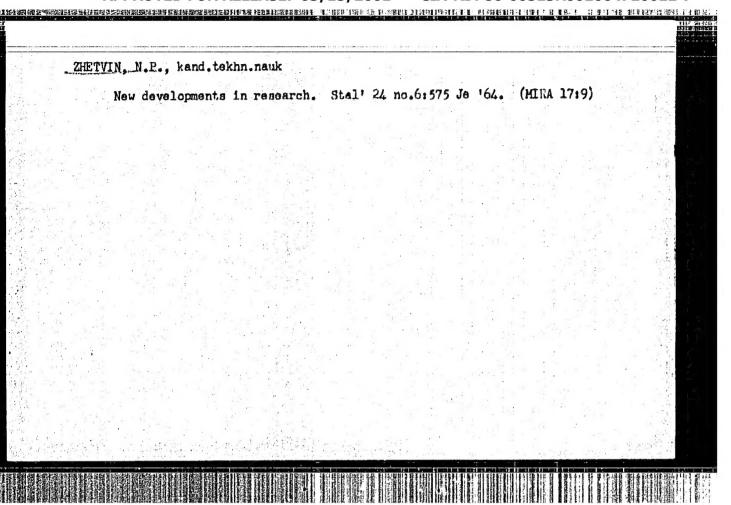
Effect of chemical composition on the properties of peoned
Khl8N9 steel. Metalloved. i term. obr. met. no.1:45-41 Ja '63.

(MIRA 16:2)

1. Zavod "Serp i molot".

(Steel alloys—Testing)



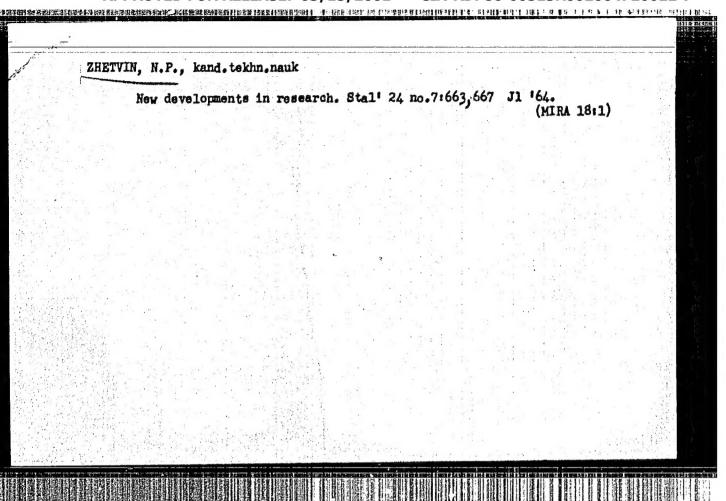


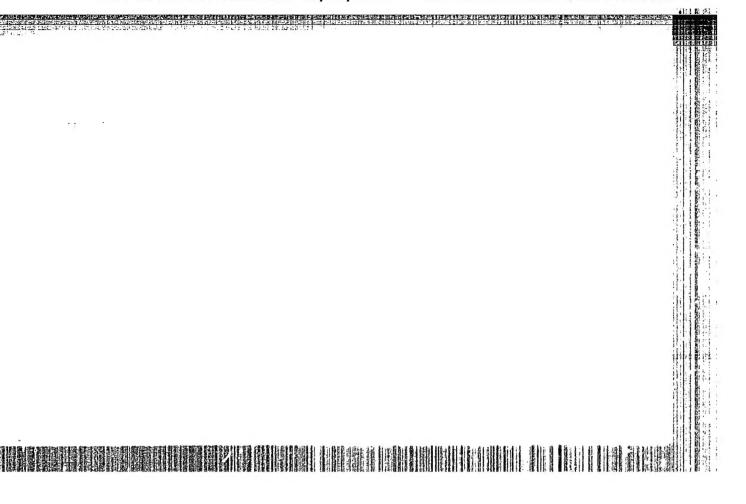
	ACC_NR: AT6034458 (A) SOURCE CODE: UR/0000/66/000/000/0213/0218	
	AUTHOR: Zhetvin. N. P.; Frid, Ya. L.; Kontsevaya, Ye. M.; Sokol, I. Ya.; Lyukovich, V. L.	
	ORG: none	
	TITIE: Study of the kinetics of hardening and softening of heat resistant alloys with the aim of choosing the temperature interval for hot plastic deformation and heat treatment	
	SOURCE: AN SSSR. Institut metallurgii. Svoystva i primeneniye zharoprochnykh splavov (Groperties and application of heat resistant alloys). Moscow, Izd-vo Nauka, 1966, 213-218	
	TOPIC TAGS: heat resistant alloy, metal deformation, metal heat treatment	
	ABSTRACT: The experiments were carried out on hot rolled samples of alloy Brand EIB28 with a thickness of 2-3 mm, and cold rolled samples of alloy Brand EF460 with a thickness of 1.0-1.5 mm. The chemical composition of the alloys is shown in the following table:	
1		
	Card 1/2	

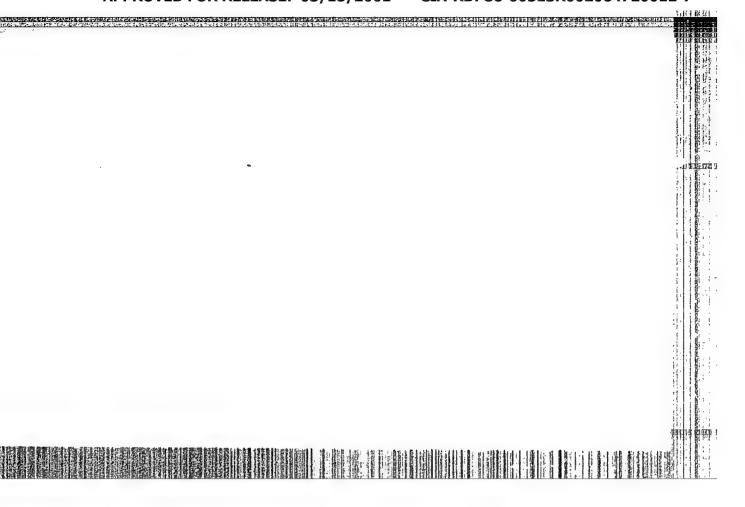
 ACC NR	AT60	34458					er- er Notes				
 Alloy EI828 EF460	• •	0,03 0,03	 Mn races races	Si 0,11 0,07	100	s 006 010	P 0,005 0,008		Ni base base	Cr 9,55 8,85	
Alloy EI828 EF460		Mo 8,8 2,2	 W 5,01	Ti 0,06 3,0	0,	B 008.	Al 4,50 1,8		Ce 0,15	Nb -	

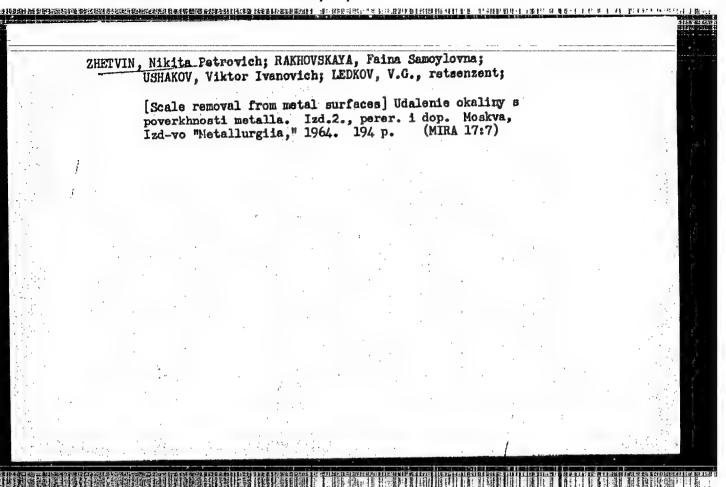
The samples were subjected to hardening in a laboratory electric furnace at a temperature of 950-1200°C, and aging at temperatures of 650-1000° with a holding time up to 12 hours. The mechanical properties (0, 6, HB, ak) and the microstructure were determined before and after aging. A phase analysis was made of the precipitates which separated out from the hardened and aged samples of alloy EIB28, and a dilatometric examination of the samples was made on a differential optical dilatometer. On the basis of the experimental data, a study was made of the kinetics and the temperature interval for the formation of the intermetallic phase of the type Ni₂Al or Ni₂(Ti, Al). The following conclusions were drawn: 1) the decomposition of the solid solutions at aging temperatures starts the minute the aging process starts; 2) a maximum degree of hardening is achieved (at 800°) in an alloy containing 27% of the intermetallic phase; 3) weakening of the aged alloy Brand EF460 is reached on heating to 1050° and above, while for alloy EIB28, this temperature is shifted to 1200°.

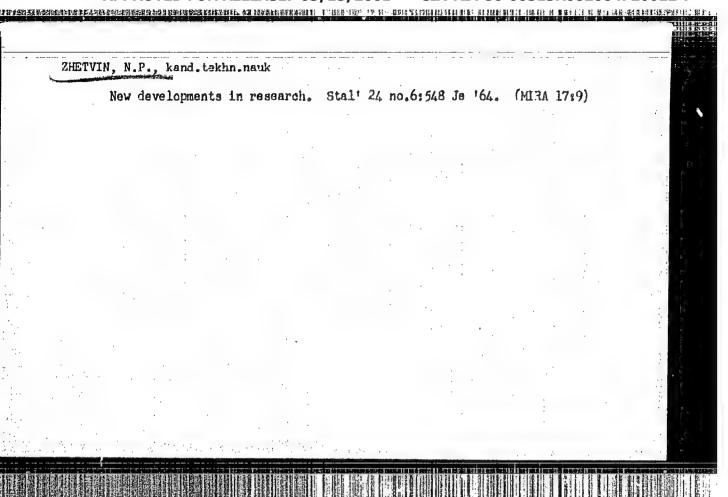
"The x ray analysis was done by S. S. Potapova, and the analysis of the intermetallic precipitate by A. P. Pogodina.". Orig. art. has: 5 figures and 2 tables.







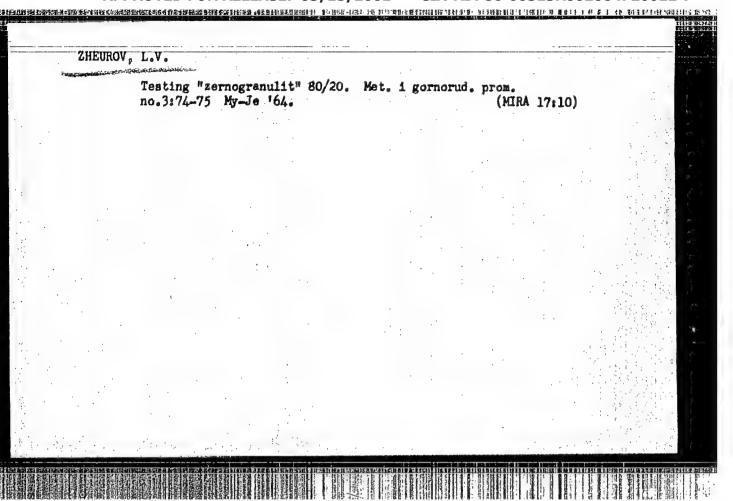




PANCHENKO, Ye.J.; PANSHINA, M.M.; REMALD, I.B.; BLINKOVA, T.M.; KRYLOVA, L.I.;
ZHDANOV, V.V.; ZHEITZA, M.P.; LUSHINS, B.G.

Remidual stresses in hillers made of ALOG steel. Stan. 1 instr.
36 no.8:27-29 Ag *65.

(MIRA 18:9)



ZHEUROV, L.V.

Work practices in using multistage bits in air drilling at the "Chikalovskii" strip mine, Met. i gornorud. prom. no.3:82 My-Je '65. (MIRA 18:11)

ZHE JAGINA, PY.

1(0) 7.3

PHASE I BOOK EXPLOITATION

sov/2835

Moscow. Aviatsionnyy institut im. Sergo Ordzhonikidze

Voprosy proyektirovaniya samoletov; sbornik statey (Problems in Aircraft Designing; Collection of Apticles) Moscow, Oborongiz, 1959. 74 p. (Series: Its: Trudy, vyp. 108) Errata slip inserted. 3,100 copies printed.

Sponsoring Agency: Ministerstvo vysshego obrazovaniya SSSR.

Ed.: A.L. Gimmel'farb, Candidate of Technical Sciences,
Docent; Ed. of Publishing House: K. I. Grigorash; Tech.
Ed.: L. A. Pukhlikov; Managing Ed.: A.S. Zaymovskaya,
Enginear.

PURPOSE: This book is intended for personnel in the design offices of aircraft plants. It may also be used by students at aviation institutes.

COVERAGE: This collection of articles describes the results of theoretical and experimental investigation connected with the determination, during the designing stage, of basic aircraft and wing parameters, total weight of aircraft and its components, type of engines and the amount of fuel. Problems of Card 1/4

Problems in Aircraft Designing (Cont.)	sov/2835
aircraft strength and stability are personalities are mentioned. Refere	also considered. No nces appear in the text.
TABLE OF CONTENTS:	
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The author determines basic parameter and selects from them the most importance: Total weight of aircraft, wing and weight, and the necessary thrus	rtant. These z-surface design
Gimmel'farb, A.L. [Candidate of Technic Calculating Necessary Fuel Supply and To Aircraft During the Designing Stage In this article the author deducts formulas based on only two static co	simple weight

Problems in Aircraft Designing (Cont.) SOV/	⁄2835
weight efficiency and fuel consumption:	
Fadeyev, N.N. [Candidate of Technical Science]. Comparative Evaluation of Aircraft Engines According to Their Weight in Flight A method is given to help in the selection of an engine for a given aircraft and for determined regimes and flight distances	41
Zhevagina, A.A. [Candidate of Technical Sciences]. Determination of Critical Stresses in Laminar Compressed Panels With Veneer Covering Results of an investigation show that sufficient support is formed for a thin veneer lining by a filling with the specific weight of 0.065 * 0.1 gr/cm ³ . With this filling the panel behaves as a homogeneous body until the moment of a general loss of rigidity.	52
Voyt, Ye.S. [Candidate of Technical Sciences]. Stability Crossed-Bar Assembly Which Has Been Compressed in One Direction The author is concerned with the plane and curved Card 3/4	of a
the first of the f	

Problems in Aircraft Designing (Cont.)

SOV/2835

reinforced panels used in ship and aircraft construction. He analyses the influence of separate factors on the stability of the panels and indicates practical methods of choosing, in the first approximation, the most convenient disposition of basic elements of the panel.

AVAILABLE: Library of Congress

Card 4/4

IS/mg 11-23-59

。A. F. A. B. 1 【"你看这些,但你你们是你没有好你的现在我们就是你们的话,你还是你们的我们的是我们的是我们的,我们们就是我们的我们的

STURMAN, A.V., veter. vrach (Strashenskiy rayon, Moldavskaya SSR); BULGAKOV, Yu.N., veter. fel'dsher (Strashenskiy rayon, Moldavskaya SSR); KAL-NITSKIY, P.I., veter. vrach (Strashenskiy rayon, Moldavskaya SSR); OCHAKOVSKIY, Z.M., veter. vrach (Strashenskiy rayon, Moldavskaya SSR); GOTSENOGA, A.D. (Strashenskiy rayon, Moldavskoy SSR); ABRAM-YAN, G.I., veter. vrach; MEKHTIYEV, M.G., veter. fel'dsher (s.Shi-rozlu, Vedinskogo rayona Armyanskoy SSR); KIRAKOSYAN, A.A., veter. vrach; GEORGIYEV, Yu.P., veter. vrach; LOMAKIN, A.M., nauchnyy sotrudnik; SHEPELEV, L.A., veter. vrach; TARASOV, I.I., assistent; ROMASHKIN, V.M., veter. tekhnik; ANDRIYAN, Ye.A.; BARTENEV, V.S.; KOROL', Ye.I., veter. tekhnik; YEROSHENKO, A.K., aspirant; BANZEN, Ya.P.; SARAYKIN, I.M., prof.; ZHEVAGIN, A.M., veter. vrach; BUT'-YANOV, D.D., veter. vrach (Klimovichskiy rayon, Mogilevskoy oblasti BSSR); SHALYGIN, B.V., veter. vrach (Klimovichskiy rayon, Mogilevskoy oblasti BSSR); RYABOKON, G.T., veter. fel'dsher; MOVSUM-ZADE, K.K., prof.; DUGIN, G.L., aspirant; TITOV, G.I., nauchnyy sotrudnik; MEDVEDEV, I.G., veter. vrach.; ALIKAYEV, V.A.; ALLENOV; O.A., veter.vrach.

Prophylaxis and treatment of noninfectious diseases in calves and piglets. Veterinariia 40 no.2:40-47 F '63. (MIRA 17:2)

1. Ul'yanovskaya oblastnaya veterinarno-bakteriologicheskaya laboratoriya (for Sturman). 2. Kolkhoz imeni Kirova. Volokonovskogo (Continued on next card)

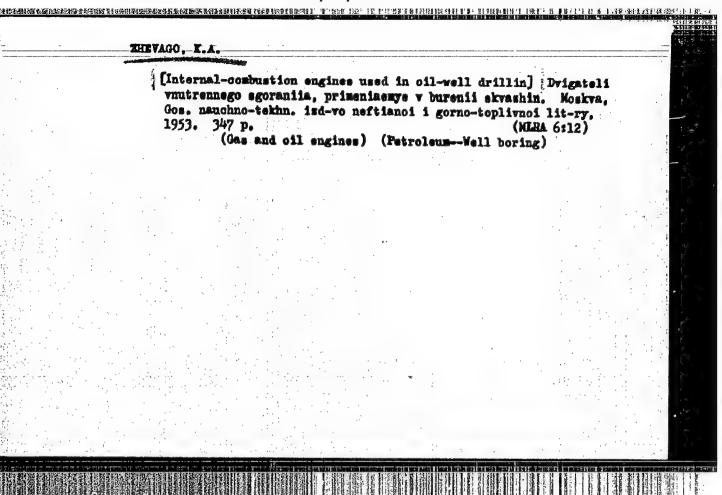
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STURMAN, A.V. (continued) Card 2. rayona, Belgorodskoy oblasti (for Bulgakov). 3. Sovkhoz "Akhuryanskiy", ArmSSR (for Abramyan). 4. El'khotovskaya veterinarno-bakteriologicheskaya laboratoriya Severo-Osetinskoy ASSR (for Allenov). 5. Shagatskiy veterinarnyy uchastok, Sisianskogo rayona, ArmSSR (for Kirakosyan). 6, Sovkhoz "Vekhno", Pskovskoy oblasti (for Georgiyev). 7. Leningradskaya lesotekhnicheskaya akademiya imeni S.M.Kirova (for Lomakin). 8. Siverskiy veterinarnyy uchastok, Gatchinskogo rayona Leningradskoy oblasti (for Shepelev). 9. Saratovskiy zooveterinarnyy institut (for Tarasov, Yeroshenko). 10. Sovkhoz "Gorodishchenskiy" Penzenskoy oblasti (for Romashkin). 11. Glavnyy veterinarnyy vrach plemennogo sovkho za imeni Litvinova, Frunzenskogo rayona, Luganskoy oblasti (for Andriyan). 12. Svinosovkhoz imeni Podtelkova, Kosharskogo rayona, Rostovskoy oblasti (for Bartenev). 13. Sovkhoz "Shakhter" Done tskoy oblasti (for Korol'). 14. Zernosovkhoz "Mikhailovskiy" TSelinnogo kraya (for Banzen). 15. Kishinevskiy sel'skokhozyaystvennyy institut (for Saraykin, Zhevagin). 16. Klimovichakiy rayon, Mogilevskoy oblasti, BSSR (for But'yanov, Shalygin). 17. Kolkhoz imeni Shevchenko Tal'novskogo rayona, Cherkasskoy oblasti, UkrSSR (for Ryabokon'). 18. Leningradskiy veterinarnyy institut (for Movsum-zade, Dugin). 19. Buryatskaya nauchno-proizvodstvennaya veterinarnaya laboratoriya (for Titov). 20. Buryatskiy seliskokhozyaystvennyy institut (for Medvedev).

ZHEVAGINA, Z. Z.

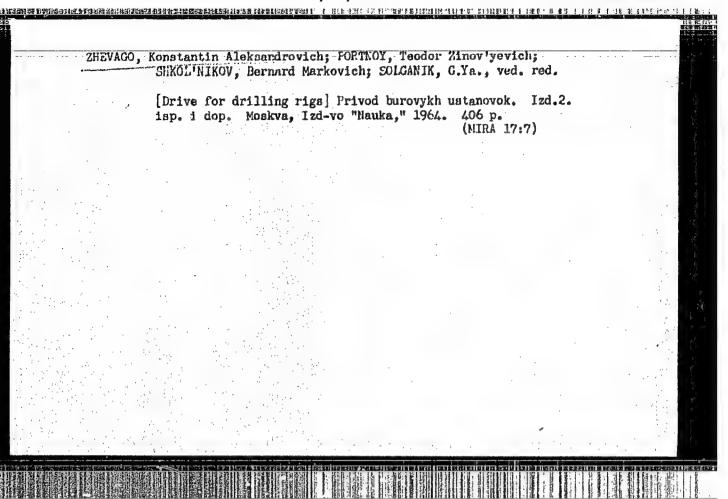
"The Determination of the Parameters of Compressed Laminated Panels Containing Filler Material." Cand Tech Sci, Moscow Order of Lenin Aviation Inst imeni Sergo Ordzhonikidze, Moscow, 1955. (KL, No 9, Feb 55)

SO: Sum. No. 631, 26 Aug 55-Survey of Scientific and Technical Dissertations Defended at USSR Higher Educational Institutions (14)



ZHEVAGO, K.A. Prinimal uchastiye MORGULIS, Yu.B.; BULATOV, S.I., red.
izd-va; EL'KIND, V.D., tekhn. red.; UVAROVA, A.F., tekhn. red.

[High-speed diesel engines; design, assembly and operation]Bystrokhodnye dizeli; ustroistvo, montazh i ekspluatatsiia. Moskva, Mashgiz, 1962. 399 p. (MIRA 16:1) (Diesel engines)



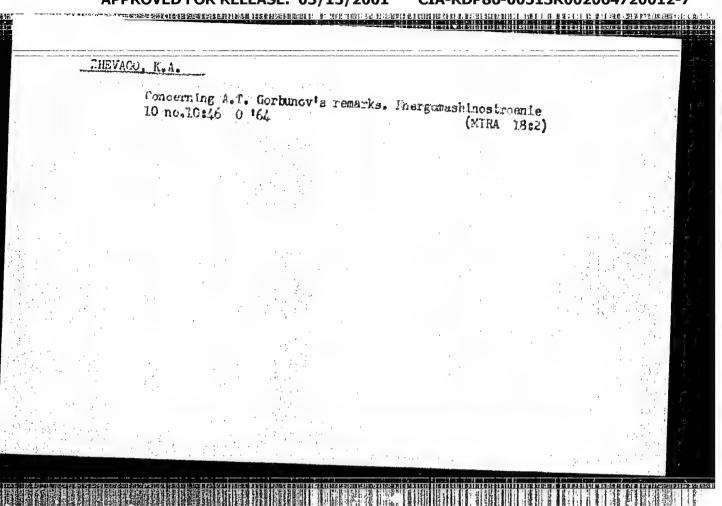
ZHEVAGO, Konstantin Aleksendrovich; BEKMAN, Yu.K., vedushchiy redaktor;

FULOSINA, A.S., tekhnicheskiy redaktor

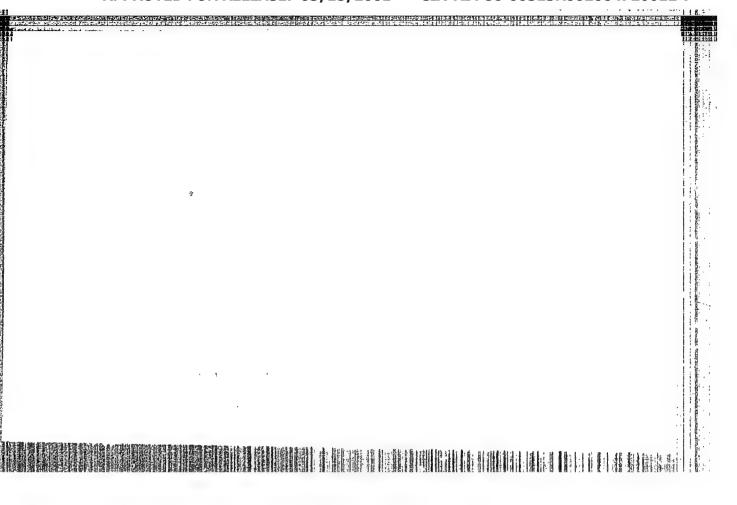
[Internal combustion engines used in well drilling] Dvigateli vnutrennego sgoraniia, primeniaemye v burenii skvashin. izd-vo neftianoi i gorno-toplivnoi lit-ry, 1956. 416 p. (MLRA 9:11)

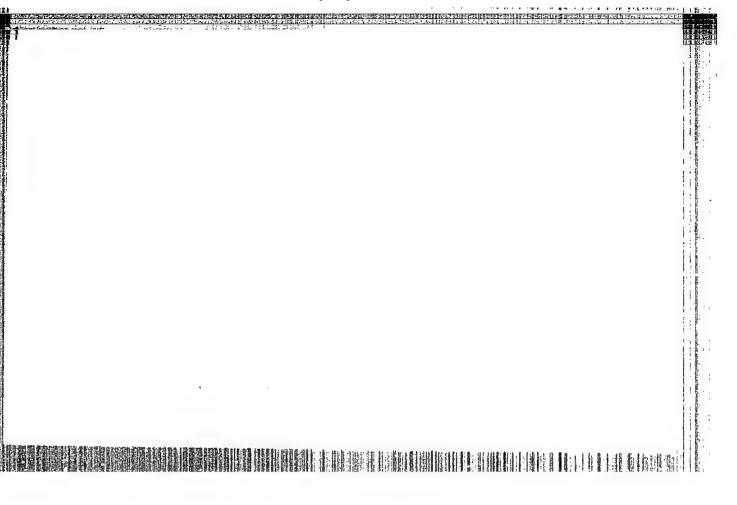
(Gas and oil engines)

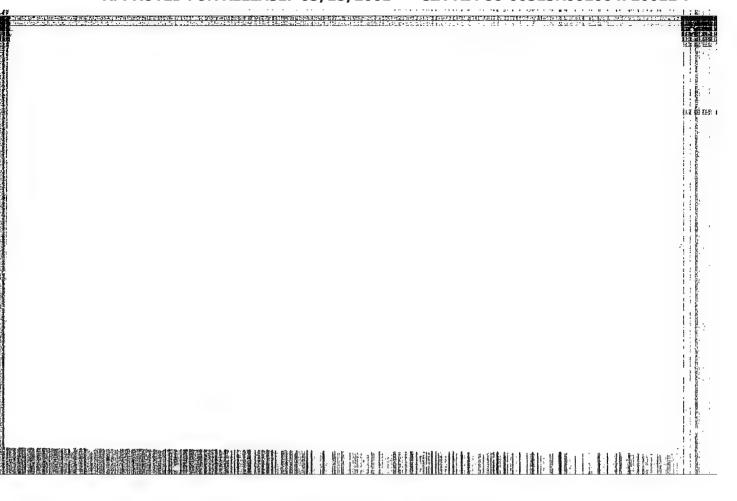
(Oil well drilling-Equipment and supplies)



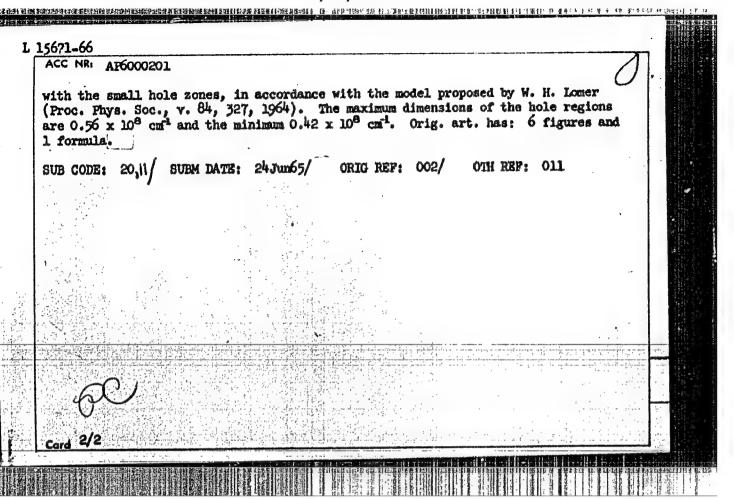
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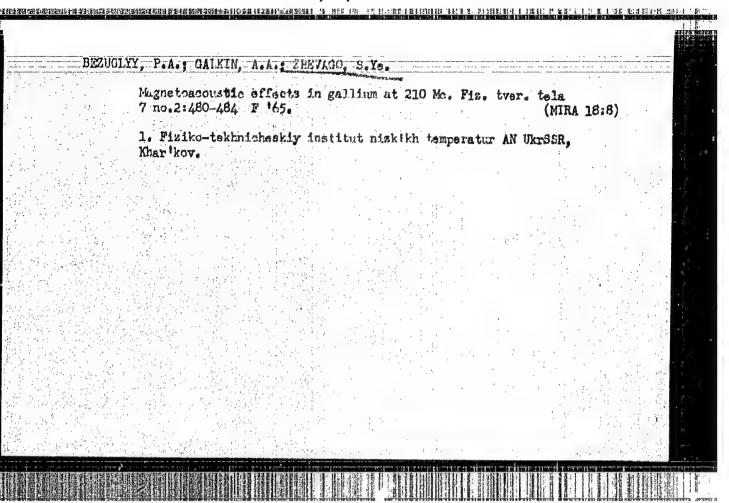


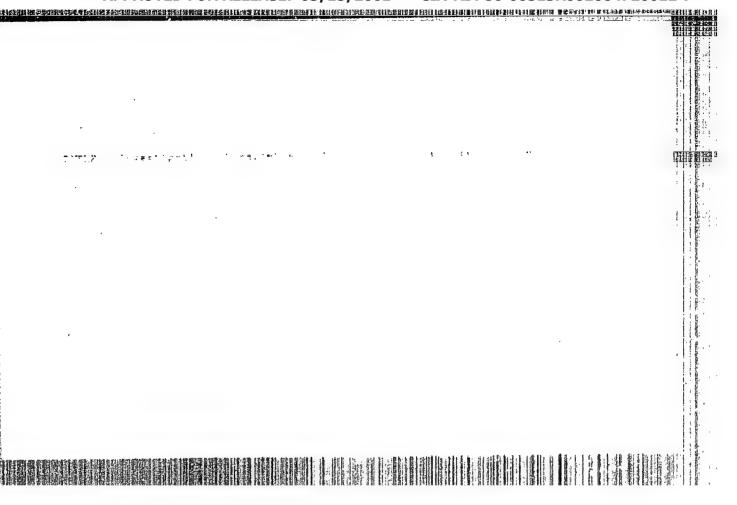


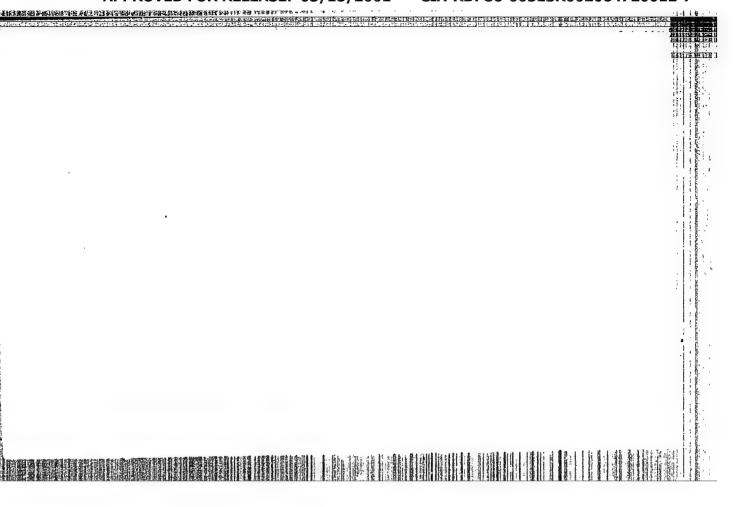


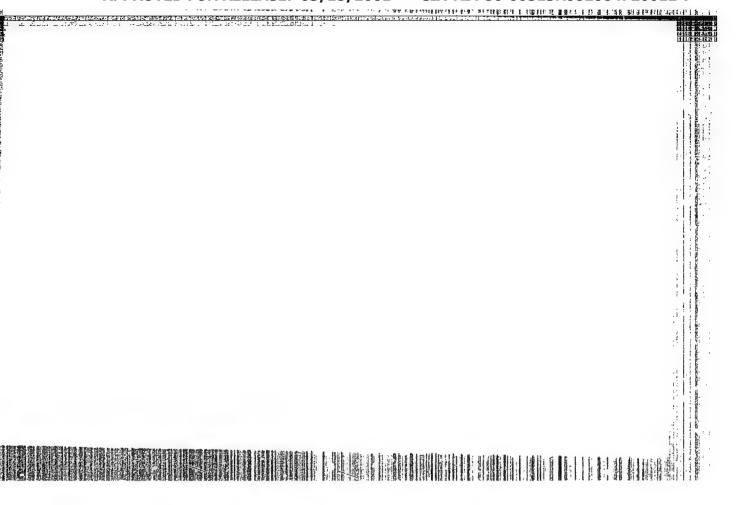
,这个人的,是一个人的,这是一个时间和时间,这个时间的时间,这个时间的时间,这个时间,这个时间,这个时间,这个时间的时间,我们可以可以的时间,我们可以可以可以可 L 15671-66 EWT(1) ACC NR: AP6000201 SOURCE CODE: UR/0056/65/049/005/1457/1462 -AUTHOR: Bezuglyy, P. A.; Zhevago, S. Ye.; Denisenko, V. I. ORG: Physicotechnical Institute of Low Temperatures, Academy of Sciences, UkrSSR (Fiziko-tekhnicheskiy institut nizkikh temperatur Akademii nauk UkrSSR) TITLE: Magnetoacoustic investigation of the Fermi surface of molybdenum SOURCE: Zhurnal eksperimental'noy i teoreticheskoy fiziki, v. 49, no. 5, 1965, 1457-1462 TOPIC TAGS: molybdenum, magnetoacoustic effect, single crystal, magnetic anisotropy, transverse magnetic field, accustic wave, electromagnetic wave explication ABSTRACT: In view of the fragmentary experimental data published so far on the magnetoacoustic effects in molybdenum, the authors investigated this effect in greater detail by studying the anisotropy of the oscillation periods of the geometrical resonance in transverse magnetic fields when the acoustic wave vector was oriented along the principal crystallographic directions of a single-crystal sample of molybdenum. The measurements at 200 Mc frequency and 4.2K used the pulse procedure of A. A. Galkin and A. P. Korolyuk (PTE, no. 6, 199, 1960). The temperatures were 1.8K in the case of q \parallel [100] and 4.2K in the case q \parallel [110] and [111]. The results showed that in different angle ranges, three different oscillation periods are observed in the absorption coefficient, one short-period and two long-period. It is shown that the short-period oscillations, observed for q | [100], give the dimensions of the electronic surface, while the long-period oscillations are associated Card 1/2.

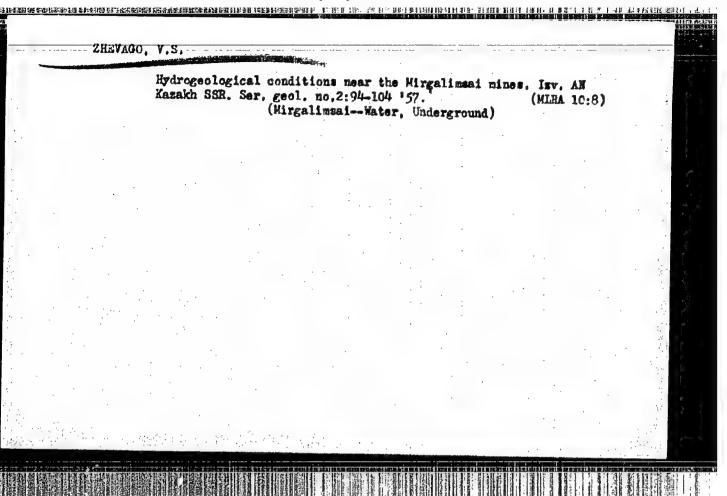


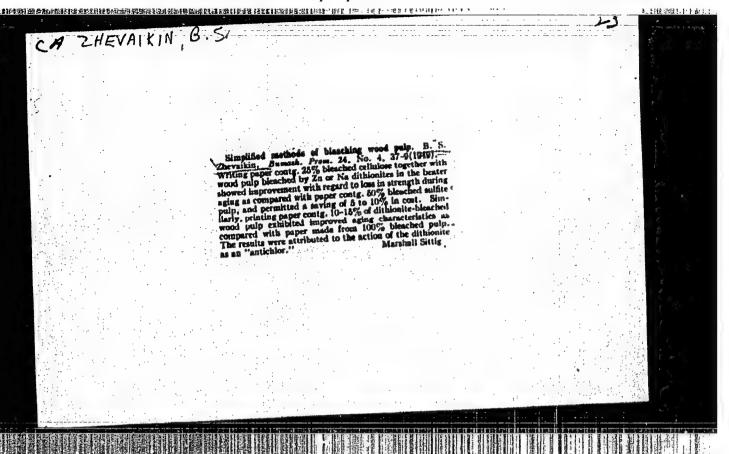












ZHEVAKHOV, D.S., dotsent, kand.tekhn.nauk Increasing the effectiveness of heat exchangers for TV-1000 mercury-ere rectifiers manufactured by the "Urnlelettroapparat" Plant. Isv.vys.uoheb.sav.; energ. 2 no.12:66-72 D '59, (MIRA 13:5) 1. Ural'skiy politekhnicheskiy institut imeni 8.M.Kirova. Predstavlena kafedroy teploenergeticheskikh ustanovok. (Heat exchangers) (Electric current rectifiers)

SOV/143-59-2-9/19 8(3) Zhevakhov, D.S., Candidate of Technical Sciences 'AUTHOR: The Investigation and Improvement of Cooling Devices TITLE: for Mercury Rectifiers (Issledovaniye i uluchsheniye okhlazhdayushchikh ustroystv rtutno-vypryamitel'nykh ustanovok) Izvestiya vysshikh uchebnykh zavedeniy - Energetika. PERIODICAL: 1959. Nr 2, pp 68-78 (USSR) The author bases his article on the experience of the ABSTRACT: plant "Uralelektroapparat" in building cooling equipment for mercury rectifiers and on experience obtained during the operation of this equipment by the Omsk RR line. For cooling mercury rectifiers, a closed primary water circulation cycle is used. heat is removed from the mercury rectifier housing by distilled water, which in turn is cooled in a heat exchange according to the system water - air. Figure 1 shows a schematic diagram of this cooling system. Initially, for cooling the mercury rectifier RMNV-500 x6, two laminar heat exchangers were Card 1/6

807/143-59-2-9/19

The Investigation and Improvement of Cooling Devices for Mercury Rectifiers

used (as in the PVS-6 ventilation system) with an axial fan Nr 7 of TsAGI and a centrifugal pump 2K6B for the distilled water. The eight-blade fan was powered by a 1.7 kw motor and had a discharge capacity of 12,000-18,000 m3/h at a pressure of 14-17 mm. The pump had a capacity of 10-25 m3/h at 16.5-22 mm water column and was powered by a 2.8 kw motor. The heat removed by this device amounted to 28,000 kcal/h. This performance was unsatisfactory to the plant designers and to the installations operating the rectifier equipment, since it did not provide the necessary cooling effect at 30°C air temperature. An investigation of the available heat exchangers showed that none of the standard types met the requirements for a rectifier cooling unit. Therefore, a new heat exchanger was designed which consisted of 96 tubes of 1 m length, having a surface of 50 m². Using the TsAGI fan, about 30,400 kcal/h of heat

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The Investigation and Improvement of Cooling Devices for Mercury Rectifiers

were removed, which is adequate for the RMNV-500x6 rectifier. The plant "Uralelektroapparat" developed on these data a heat exchanger with riffled tubes which was mass-produced later on. At a large number of installations, these cooling units worked satisfactorily, with exception at the Omsk RR, where an air temperature of +40°C and a humidity of 53% were observed. In addition the rectifiers had to work constantly on full capacity. The heat to be removed amounted to 30,000 kcal/h. The series cooling units were therefore inadequate. For this reason experiments were conducted for increasing the heat transfer by spraying the heat exchanger with water. Figure 2 shows the so-called irrigation cooling. It was established that the irrigation cooling increased the heat transfer factor of a riffled surface by at least three times. However, the plant "Uralelektroapparat" was not satisfied with this method, since

Card 3/6

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The Investigation and Improvement of Cooling Devices for Mercury Rectifiers

the designers feared corrosion effects, rusting, etc. The irrigation cooling proved its effectiveness at a number of RR substations. The plant developed a new type of heat exchanger consisting of oval tubes, as shown by figure 5. Table 2 shows a comparison between the round, riffled tubes and the oval ones, whereby the latter require less space. The difficulties in producing and fastening of the oval tubes have been overcome recently and will not be an obstacle for the mass-production of this new heat exchanger. However, the application of the oval tubes does not solve problems as they occur during summer at the Omsk RR. Therefore various other cooling methods were investigated: 1) using a water - water heat exchanger with a cooling tower; 2) using a Freon refrigerator; and 3) using an ice tank. In his conclusion, the author points out that the refrigeration equipment should not be used,

Card 4/6

SOV/143-59-2-9/19

The Investigation and Improvement of Cooling Devices for Mercury Rectifiers

基**在14月24日的**中社的特殊基础的经验基础的逐渐和经济基础和基础的基础的基础的数据转要。这是14日的主题的主题中一个时间主题,可是

since the AK-2-FU-30/15 refrigerator unit costs twice as much as a water cooling system with a cooling tower. Ice may be used as a cooling agent only in those areas where it can be readily produced. Experiments conducted in 1936 by Engineer Gogolin showed that about 167 kg/h of ice were required for obtaining a cooling efficiency of 3,000 kcal/h. The author recommends the application of ice tanks for West Siberia, where a cooling unit would be used only during one month in summer. Finally, the author recommends moistening the heat exchanger surfaces on the condition that the latter are protected by a corrosion-resistant coating. There are 5 diagrams, 2 graphs, 2 tables and 1 Soviet reference.

ASSOCIATION:

Ural'skiy politekhnicheskiy institut imeni S.M. Kirova (Ural Polytechnical Institute imeni S.M. Kirov)

Card 5/6

SOV/143-59-2-9/19

The Investigation and Improvement of Cooling Devices for Mercury Rectifiers

Kafedra teploenergeticheskikh ustanovok of Thermal Equipment) (Chair PRESENTED:

August 31, 1958 SUBMITTED:

Card 6/6

DESCRIPTION OF THE PROPERTY OF

8(6) 50V/143-59-12-9/21

AUTHOR: Zhevakhov, D.S. Docent, Candidate of Technical Sciences

TITLE: Raising the Efficiency of TV-1000 Heat Exchangers of

Mercury Rectifiers from the "Uralelektroapparat" Plant

PERIODICAL: Izvestiya vysshikh uchebnykh zavedeniy: Energetika,

1959, Nr 12, pp 66-72 (USSR)

ABSTRACT: This is an account of the TV-1000 heat exchanger, how

it works, and its performance figures. It was improved in design by the Kafedra teploenergeticheskikh ustanovok Ural'skogo politekhnicheskogo instituta (Chair of Thermal Power Installations of the Urals Polytechnical Institute) in collaboration with the "Uralelektroapparat" Zavod ("Uralelektroapparat" Plant), and is designed for the circulation cooling of mercury rectifiers on a rectified current of 1000 a by means of distilled water introduced into the cavity of the rectifier: this water is it-

self cooled by water from the source of supply. Figure 1

Card 1/4

SOV/143-59-12-9/21

Raising the Efficiency of TV-1000 Heat Exchangers of Mercury Rectifiers from the "Uralelektroapparat" Plant

is a photograph of the TV-1000 heat exchanger, Figure 2a is a diagram of it in its original form and Figure 2b shows it after reconstruction. Table 2 gives performance figures for the heat exchanger before and after the reconstruction. Before, the coefficient for heat exchange was 886, afterwards - 1710 and 3530.

Its original drawbacks were that it had only one water passage, the water moved comparatively slowly and the cooling surface was limited. To obviate these, a system of ring pipes was adopted, which increased the heating surface, although the size of the apparatus remained the same. It was very important to raise the efficiency of the heat exchanger without making basic alterations in its production process. Figure 3 shows the altered construction of one of the pipe couplings of the water chambers: nr. 1 indicates the new pipe plate into which 8/10 mm pipes are rolled, placed inside 17/19 mm pipes; 2 - an additional water chamber; 3 - an additional in-

Card 2/4

SOV/143-59-12-9/21

Raising the Efficiency of TV-1000 Heat Exchangers of Mercury Rectifiers from the "Uralelektroapparat" Plant

let pipe. Before reconstruction, distilled water from the mercury rectifier entered the space between the ring pipes from inlet 1; after the reconstruction it came through inlet 3 and reached the 8/10 mm pipes. Figure 4 shows the experimental stand. Test conditions corresponded to those for normal operations. Table 1 shows the results. The method put forward by the Czech engineer Schneller / Ref. 1 / for comparing the various cooling systems was used. Figure 5 gives the comparative results. The two methods of introducing the water are approximately equal in effect. The relationship of the heat exchange coefficient to consumption of cooling water is shown in Figure 6. A specimen model of the new heat exchanger has been produced at the plant for a mercury rectifier on a rectified current of 10,000 a. It consists of 3 elements, instead of the former 10,

Card 3/4

SOV/143-59-12-9/21

Raising the Efficiency of TV-1000 Heat Exchangers of Mercury Rectifiers from the "Uralelektroapparat" Plant

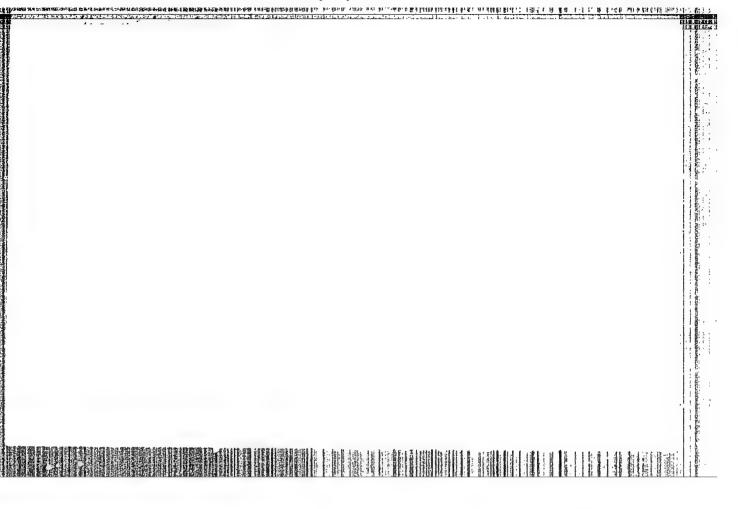
and is now being tested and perfected. The proposed reconstruction method will greatly increase the efficiency of operating installations and this may be specially important for rectifier substations on electric railways. There are 2 photographs, 3 diagrams, 1 graph, 2 tables and 1 Czech reference.

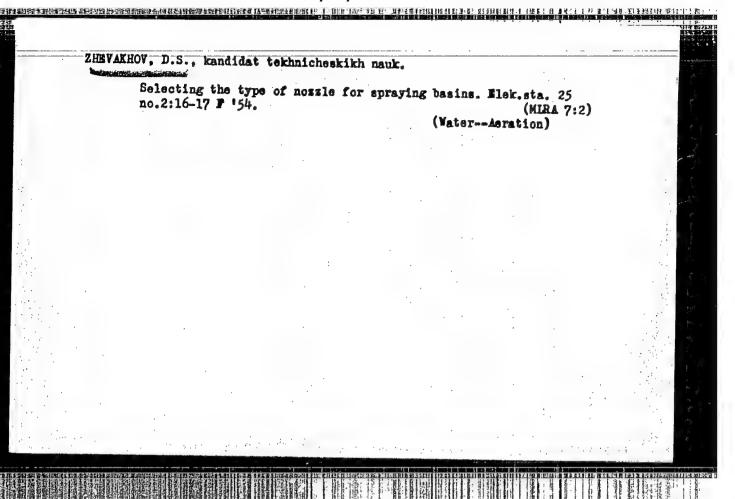
ASSOCIATION: Ural'skiy politekhnicheskiy institut imeni S.M. Kirova (Urals Polytechnical Institute imeni S.M. Kirov)

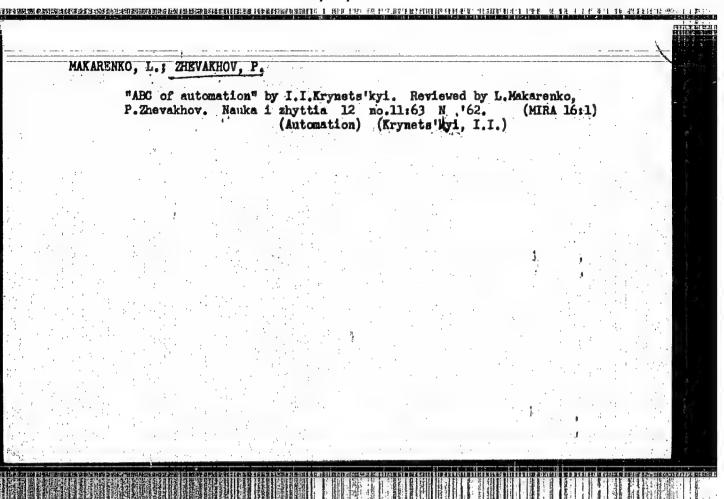
SUBMITTED: June 27, 1959, by the Kafedra teploenergeticheskikh

ustanovok (Chair of Thermal Power Installations).

Card 4/4







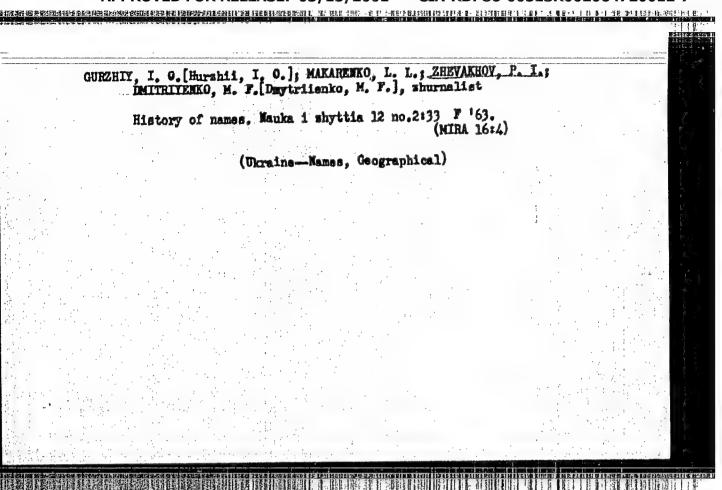
GURSHIY, I.O. [Hurzhii, I.O.], doktor isotr.hank; MAKARENKO, L.L.; ZHEVAKHOV, R.I.; DMYTRIYENKO, M.F. [Dmytriienko, M.F.], zhurnalist

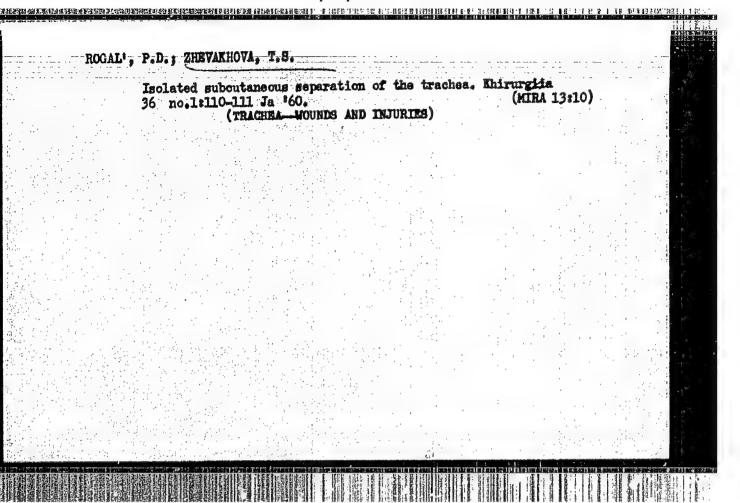
History of names. Nauka i zhyttia 12 no.1:17 Ja '63. (MIRA 16:3)

1. Chlen-korrespondent AN UkrSSR (for Curzhiy). 2. Direktor Gosudarstvennyy istoricheskoy biblioteki UkrSSR (for Makarenko).

3. Clavnyy bibliotekar' Gosudarstvennoy istoricheskoy biblioteki UkrSSR (for Zhevakhov).

(Donets Basin-Names, Geographical)





ZHEVAKHOVA, T.S., zasluzhennyy vrach UkrSSR (Cherkassy (obl.), ul. Shevchenko, dom 242)

Ossifying hematoma of the elbow joint. Ortop., travm. i protez. no.8:65-66 '62. (MIRA 17:10)

ZHEVARHOVA, T.S., zasluzbennyy vrach USSE

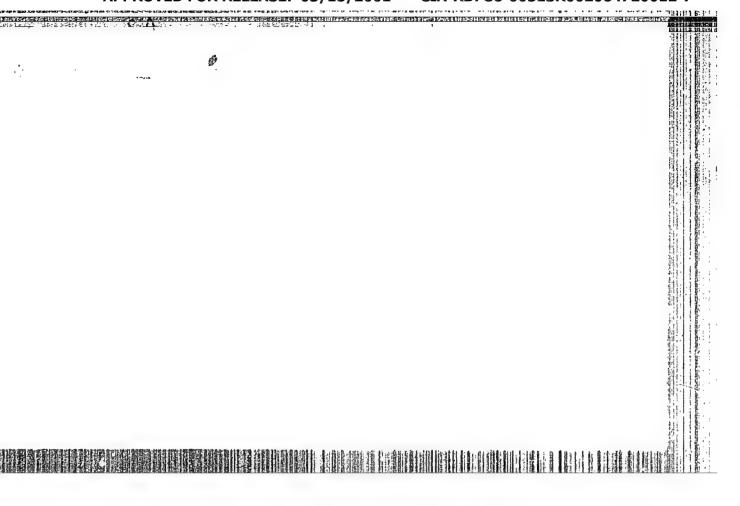
Treatment of epicondylitis. Ortop., travm.i protes. 20 no.12:
57 D'59. (MEA 13:5)

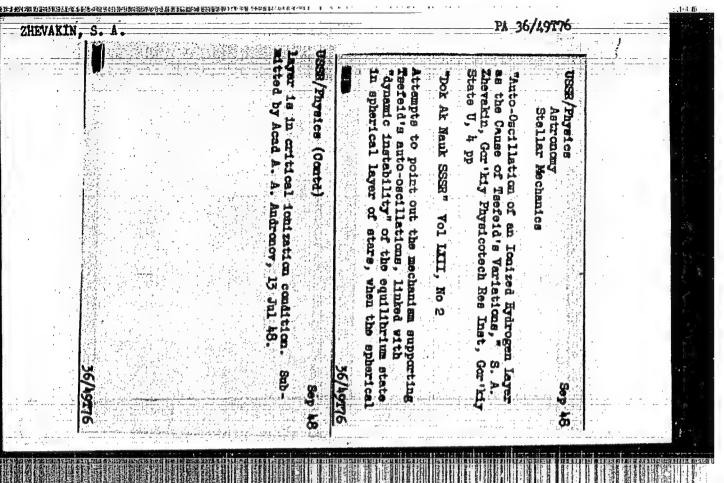
1. Iz ortopedicheskogo otdeleniya Cherkanskoy gorodskoy bol'nitsy.
(HUMRHUS--DISEASES)

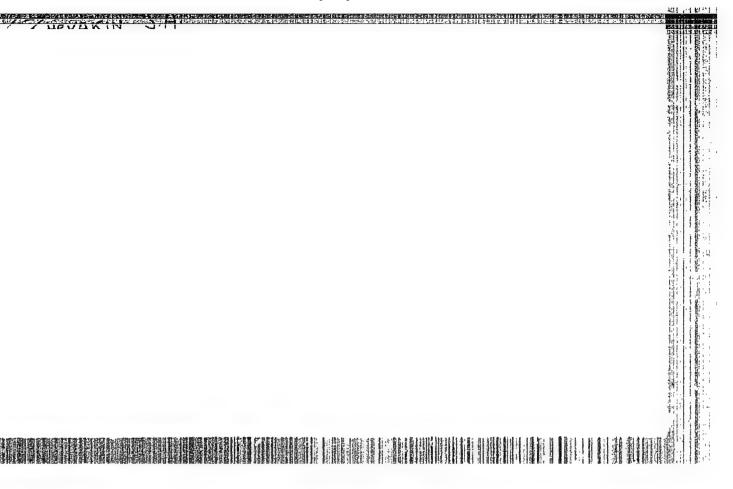
ZHEVAGO, Konstantin Aleksandrovich; PORTNOY, Teodor Zinov'yevich;
SHKOL'NIKOT, Bernard Markovich. Prinimal uchastiye SUD, I.I..
MARTINOYA, M.P., vedushohiy red.; POLOSINA, A.S., tekhn.red.

[Boring equipment drives] Privod burovykh ustanovok. Moskva,
Gos.nauchno-tekhn.isd-vo neft. 1 gorno-toplivnoi lit-ry, 1960.

362 p. (Boring machinery)







ZHEVAKIN, S. A.

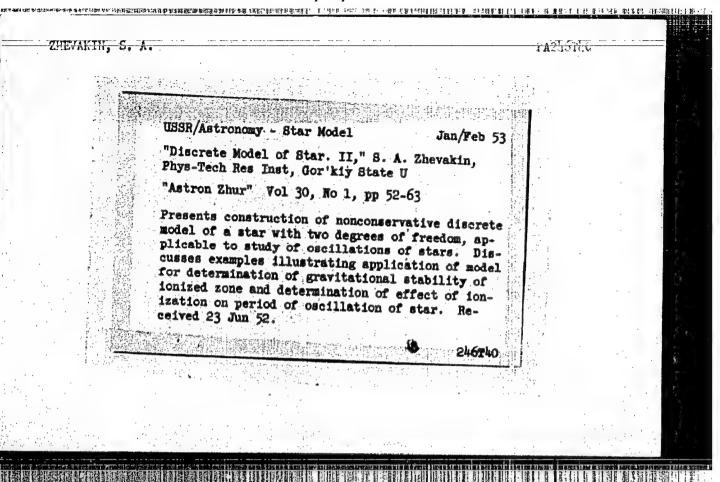
USBR/Astronomy - Stellar Fluctuations Jan/Feb 52

"Discrete Model of Star (Oscillations and Explosion). I," S. A. Zhevakin, Gor'kiy Phys-Tech Res Inst

"Astron Zhur" Vol XXIX, No 1, pp 37-48

Shows that properties of such model are similar to those of models available in astrophys literature. Studies auto-oscillations of model with one deg of freedom, sustained by action of swinging mechanism connected to thermonuclear energy release. Shows that nuclear reactions are not the cause of stellar oscillations. Received 15 Feb 51.

20215



- 1. ZHEVAKIN. S. A.
- 2. USSR (600)
- 4. Stars, Variable
- 7. On the theory of the Cepheids. Part 1, Astron. zhur., 30, No. 2, 1953.

Attempts to prove that variability of Cephoids is due to a zone of doubly ionized H_c on their periphery and therefore knowledge of internal structure of star is not essential. Theory of Cepheids may be helpful for classification of internal structure of variables. Received 23 Jun 52.

9. Monthly List of Russian Accessions, Library of Congress, April, 1953, Uncl.

AID - P-231

ning, mil n. V.

Subject

USSR/Astronomy

1/2

Author

Card

Zhevakin, S. A.

Title

On the Theory of Stellar Changeability. II

(Phase Displacement between the Variations in Brightness and in Radial Velocity in the Cepheid and Long-period

Variables)

Periodica1

Astron. zhur., v. 31, 2, 141-153, Mr - Ap 1954

Abstract

Unsoundness of known present-day attempts to explain the phase displacement between the variations in radial velocity and in brightness of variable stars is demonstrated.

A multilayer discrete model of a fluctuating non-adiabatic envelope of a star is constructed. On the basis of this model a theory is developed of non-adiabatic pulsations of the envelope. The theory is of interest: 1) in the study of auto-variations of variable stars of "large sequence", the variations being sustained by the negative dissipation created in the zone of critical

ionization He II; and 2) for the study of the conditions

Astron. zhur., v. 31, 2, 141-153, Mr - Ap 1954, (additional card)

AID - P-231

Card

2/2

of the origin of the phase displacement between the variations in radial velocity and in the brightness, characteristic of the cepheids and long-period variables. Formulae. 23 references (after 1938),

Institution:

Gor'kov Physico-Technical Institute, Gor'kov University

Submitted

July 5, 1953

ZHEVAKIN, S.A.

Subject

USSR/Astronomy

AID P - 427

Card 1/2

Pub. 8, 6/16

Author

: Zhevakin, S. A.

Title

On the Theory of Stellar Changeability. III (Phase Displacement between the Variations in Brightness and In Radial Velocity in the Cepheids and Long-period Variables)

Periodical

: Astron. zhur., v. 31-4, 335-357, J1-Ag 1954

Abstract

that under certain conditions a phase lag between the maximum emission of radiation and the phase of maximum contraction of a star is equivalent to 1 of the oscillation period. Simultaneously negative dissipations author establishes that in a single layer atmosphere change the phase displacement in the cepheids. Formulae, 10 tables, 5 graphs, 7 references.

Astron. zhur., v. 31-4, 335-357, J1-Ag 1954

Card 2/2

Pub. 8, 6/16

Institution: Gor'kiy Physico-Technical Institute, Gor'kiy University

Submitted

: July 5, 1953

CIA-RDP86-00513R002064720012-7" APPROVED FOR RELEASE: 03/15/2001

USSR/ Physics Radio wave propagation Card : 1/1 Pub. 118 - 7/7 Authors : Getmantsev, G. G., Zhevakin, S. A., Kobrin, H. M., and Hiller, M. A. Title ! Propagation of radio waves Periodical: | Usp. fiz. nank 53/2, 298 - 303, June 1954 : The book "Propagation of Radio Waves", written by V. N. Kessenikh, is criticized. Many fundamental errors in interpretation of the subject Abstract covered by the book were found. Also, the unmethodical arrangement of many experimental data, included in the book, render it useless even for reference. In short, the publication of the book by the "Gostekhizdat" (State Publ. Rouse for Tech. Literature) is considered Institution : Submitted

"你明,""明,那中心的答案" 治治 利益 医双重磁带的过程电流电影工程电影 医毛毛性针 制作 化水压物 电光 排 电探点工法 解 用的 经经济保护服务的 经经济特别的 计 A HEVAKIN, SA USSR/Astronomy - & Cephei Card 1/1 Pub. 22 - 8/40 Authors Zhevakin, S. A. Title Regarding the theory of star variability Periodical : Dok. AN SSSR 99/2, 217-220, Nov 11, 1954 Abstract : A theory on the variability of the I and II Class (after Baade) stars (§ Cephei type) is described. The theory adequately explains almost all phenomena observed. It is based on the assumption of a radial self-oscillation of stars due to the so-called "negative dissination" (of energy, which occurs in critical zones of helium II ionization and which is return is due to the hightemmerature of the zones T(= 35,000-55,0000). Five references, 3-MSSR (1933-Institution: Physico-Technical Institute of the State University at Gorkiy Presented by: Academician M. A. Leontovich, June 16, 1954

USSR/ Astronomy - Variable stars

Card 1/1 Pub. 22 - 4/40

Authors : Zhevakin, S.A.

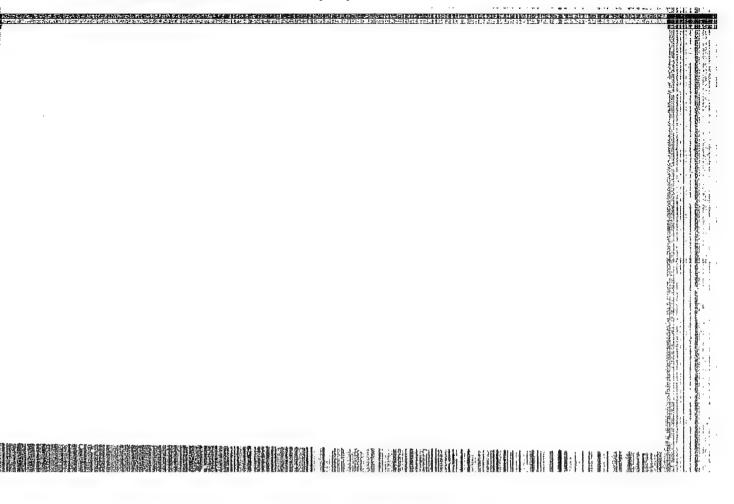
Title : On phase lagging in glitter variations (vibrations) with respect to those in energy radical-velocity

Periodical : Dok. AN SSSR 99/3, 353-356, Nov 21, 1954

The phase lagging in glitter variations with respect to those in energy radical-velocity of variable stars is explained. The explanation is substitution of a sample of exemplary data obtained through actual observation of 3 Cephes. Five references; 3-USSR (1947-1954). Piagrams;

Institution: Gorki State University

Presented by: Academician M.A. Leontovich, June 16, 1954



"APPROVED FOR RELEASE: 03/15/2001 CIA-RDP86-00513R002064720012-7 的基础的数据或数据重要的数据的数据重要数据处理数 网络亚维加强加强格利亚 一定:这样的一个银行,可是一种发展,在第三人和自然经验和自然经验和自然的。

MENUKING O'H

USSK / Radio Physics. Propagation of Radio Waves.

Abs Jour : Ref Zhur - Fizika No 3, 1957, No 7318

Author

: Zhevakin, S.A., Kobrin, M.M.

Inst

: Gor'kov University

Title

: Calculation of the Field Intensity of the Sky Wave at Short Ra-

dial Waves.

Orig Pub : Uoh. zap. Gor'kovsk, un-ta, 1956, 30, 92-136

Abstract : A method is proposed for calculating the field intensity of the space wave in the case of ionospheric propagation; this method takes into account separately the absorption in various ionizing layers. Calculations are given for the absorption in the D layer and for the case of reflection from the E, F1, and F2 layer. Equations for the absorption coefficient, obtained by using a model of a simple layer, make it possible to calculate the field intensity from specified radiation and propagation conditions (transmitter power, directivity pattern and impedance of the antenna, wavelength, zenith angle of the sung, and the length and

Card

: 1/2

APPROVED FOR RELEASE: 03/15/2001 CIA-RDP86-00513R002064720012

USSR' / Radio Physics, Propagation of Radio Waves.

I-6

Abs Jour : Ref Zhur - Fizika No 3, 1957, No 7318

Abstract : direction of the communication route) using the quantity for = fere cos 4 / (if or is the critical frequency of the E layer), which depends on the solar activity, and using the universal constants $B_{\overline{D}}$ and $A_{\overline{E}}$, determined from the field-comparison results. The equations obtained were experimentally confirmed in the 25 -- 120 meter band. Empirical equations suitable for the field intensity during night time propagation, as well as many nomograms to facilitate the calculations, are also given. Bibliography, 36 titles.

Card

: 2/2

Eneverin, 3 A.

USSR/Radiophysics - Radio-trave Propagation. Ionosphere. I-6

Abst Journal: Referat Zhur - Fizika, No 12, 1956, 35298

Author: Zhevakin, S. A., Fayn, V. M.

Institution: None

Title: On the Theory of Nonlinear Effects in the Ionosphere

Original

Periodical: Zh. eksperim. i teor. fiziki, 1956, 30, No 3, 518-527

Abstract:

In the calculation of the nonlinear effects in the ionosphere, the authors use a velocity distribution function for the electrons, obtained by one of the authors (Referat Zhur - Fizika, 1956, 1313) for the case of propagation of an amplitude-modulated high frequency field of arbitrary amplitude B_0 , in the presence of a permanent magnetic field. This makes it possible to calculate the values of the cross-modulation and other nonlinear ionospheric effects without assuming the magnetic field of the wave up be small, as cans done earlier by other authors. It is shown that even at transmitter powers greater than 250 kw and under usual conditions of radiation and propagation of

Card 1/2

USSR/Radiophysics - Radio-wave Propagation. Ionosphere, I-6

Abst Journal: Referat Zhur - Fizika, No 12, 1956, 35298

Abstract: radio waves, a noticeable deviation occurs from the results of the approximate theory of cross-modulation (linear with respect to the square of the tude of the field Eo of a strong station). Thus, in the example under consideration, at a transmitter power of 500 kg, the factor characterizing the depth of the cross-modulation, assuming collisions between the electrons and molecules, is calculated from the exact theory to be 0.465, but the linear approximation (with respect to E_0^2) results in \$3350 assuming collision with ions, this factor becomes 0.455 and 0.056 respectively. A calculation is made of the nonlinear effect of phase self-modulation, occurring upon the passage through the ionosphere of an amplitude modulated radio-wave. It is shown that this effect amounts to several radians per second, i.e., it can be detected experimentally, and used to study the ionosphere. Bibliography, 10 titles.

ZHEVAKIN, S.A. Energy transfer in a medium of nonhomogeneous temperature with turbulent convection. Astron. Zhur. 33 no.2:137-143 Nr-Ap '56. (MURA 9:8) 1. Fisiko-tekhnicheskiy institut Gor'kovskogo gosudarstvennogo universiteta. (Heat--Convection)

704

PHASE I BOOK EXPLOITATION

Zhevakin, S. A.

Teoriya pul'satsionnoy zvezdnoy peremennosti; avtoreferat dissertatsii na soiskaniye uchenoy stepeni doktora fiziko-matematicheskikh nauk (Pulsation Theory of Stellar Variability; Abstract of a Dissertation Offered for the Degree of Doctor of Physical and Mathematical Sciences) Leningrad, 1957. 8 p.

Sponsoring Agency: Leningradskiy ordena Lenina gosudarstvennyy universitet imeni A.A. Zhdanova.

PURPOSE: This booklet is the author's abstract of his dissertation for the degree of Doctor of Physical and Mathematical Sciences and is intended for advanced students of astronomy and professional

COVERAGE: The object of the dissertation is to present a theory of pulsation variability for variable stars of "large sequence"

Card 1/5

population I and "large sequence" population II (Baade). Each of these sequences is characterized by its own correlations of the spectrum and luminosity periods. The existence of these correlations which relate variable stars of the Cepheid type, semi-regular and irregular variables, and the long-period variables entering into each of the sequences leads to the postcommon nature of variability of these star types. ulate of a The theory assumes that the helium montent in the atmospheres of variable stars is 10 to 15 percent by the number of atoms. The theory leads to several conclusions concering the helium content in atmospheres of variable stars, the masses of variable stars, the degree of concentration of mass of the variable star at its center, the role of convection in transferring energy within the ionized zone of the variable star, etc. The author states that his study represents a first attempt to apply this theory in this direction and by no means exhausts all possibilities.

Card 2/5

The dissertation consists of 7 chapters, each of which the abstract annotates. Ch. I., the introduction, deals with the physical content of the theory of stellar variability and the principal results to which it leads. In Ch. II an analysis is given of ordinary oscillations of a distributed stellar model in a quasical approximation. Dissipation of energy of the star's own ordinary oscillations—is determined by this approximation. Ch. III diagusses the non-adiabatic oscillations of the ionized and the stellar atmosphere above it as observed with the selected with the view of getting—its quasi-adiabatic oscillations as close as possible to the quasi-adiabatic oscillations of the distributed model analyzed in Ch. II. It is indicated that the use of an ordinary single-layer discrete model brings qualitatively correct results. In Ch. IV. the non-adiabatic oscillations of the model, which makes more precise the results obtained in Ch. III.

Card 3/5

The next chapter indicates that the transfer of energy on the envelope in variable stars must take place by radiation and not by convection. Ch. VI deals with non-adiabatic oscillations of the stellar model. The similarity of ordinary non-adiabatic oscillations of a distributed model to the corresponding adiabatic oscillations of the same model testify to the fact that ordinary oscillations may be studied on a discrete model, so long as the latter's nonadiabatic oscillations remain close to the adiabatic oscillations of the same discrete model. This supports the method of discrete treatment of non-adiabatic stellar oscillations, which forms the basis of this dissertation. The error of D.A. Frank-Kamenetskiy on non-adiabatic stellar oscillations and the Cepheid theory is also pointed out. The last chapter refutes the objections of D.A. Frank-Kamenetskiy to the theory of stellar variability presented in the dissertation. The objections of L.E. Gurevich and A.I. Lebedinskiy are also reviewed and traced to a misunder-standing. The conclusions of the author's dissertation have

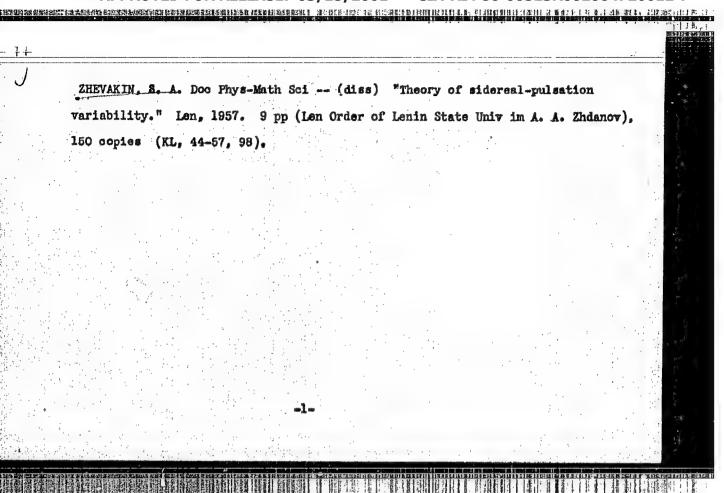
Card 4/5

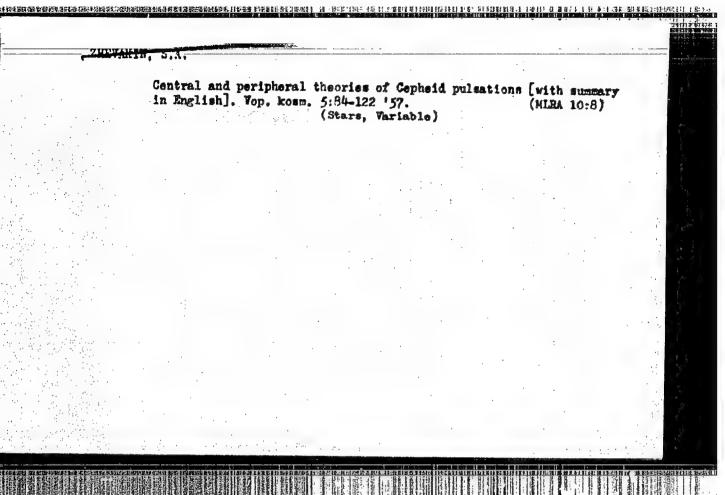
already appeared in articles published in the Astronomichesky Zhurnal, 1953, 1954, and 1955; Doklady Akademii nauk, 1954; Sbornik pamyati akad. A.A. Andronova, AN SSSR, 1955; Trudy no references.

AVAILABLE: Library of Congress (QB835.Z48)

BK/ksv 11-22-58

Card 5/5





3(1) SOY/33-35-4-7/25 AUTHOR: Zhevakin, S.A. The Dissipation of the Energy of Oscillation of a Pulsating TITLE: Star (O dissipatsii energii kolebaniy pul'siruyushchey zvezdy) PERIODICAL: Astronomicheskiy shurnal, 1958, Vol 35, Nr 4, pp 583-596(USSR) ABSTRACT: The author deduces an approximative expression for the energy dissipation of a radial, fundamental quasidiabatic oscillation of a star under radiative transfer of energy. The error of the obtained expression is shown to be at most 4%. The formula of the author gives informations on the parameters on which the dissipation of oscillation energy actually depends, provided that the concentration of mass of the star

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for the theory of stellar variability.

towards the center is sufficiently high. It is of interest

APPROVED FOR RELEASE: 03/15/2001

124-58-9-9570

Translation from: Referativnyy zhurnal, Mekhanika, 1958, Nr 9, p 16 (USSR)

AUTHOR:

Zhevakin, S. A.

新記錄後後外**分表和共享發展。新聞新聞的日報本語(1930年**8月 1811年),1911年1月1日,1921年(1931年)

TITLE:

On the Self-excited Oscillations of Variable Stars of the "Great Sequence" (Ob avtokolebaniyakh peremennykh zvezd "bol shoy posledovatel nosti")

PERIODICAL: V sb.: Pamyati Aleksandra Aleksandrovicha Andronova.

Moscow, Izd-vo AN SSSR, pp 629=656

ABSTRACT:

A summary of the results of preceding works of the author (ref. Astronom. zh., 1953, Vol 30, Nr 2, pp 161-179; 1954, Vol 31, Nr 2, pp 141-153; Nr 4, pp 335-337) on the perturbation theory of the oscillation of stars as applied to the He II ionization zone. Assuming that the shell of a star is constructed according to a polytrope with index 3, the values of the radius and the mass that satisfy the theory are sought. For 6 Cephei a radiant flux $L=7.9 \times 10^{30}$ erg/sec is obtained for $R_0 = 48R_{\odot} = 3.36 \times 10^{12}$ cm and $M=6.75M_{\odot} = 1.35 \times 10^{34}$ g. The revised value of the mass is significantly lower than the value that corresponds to the mass/radiant-flux ratio. The values obtained for the radius agree well with Wesselink's

Card 1/2

124-58-9-9570

On the Self-excited Oscillations of Variable Stars (cont.)

results (Wesselink, A., Bull. Astron. Inst. Netherl., 1947, Vol 10, Nr 380); by contrast, the results of the photoelectric measurements by Stebbins (Stebbins, J., Astrophys. J., 1945, Vol 101, Nr 1) cannot be reconciled with theory. The calculation is extremely sensitive to the assumed value of the index of the polytrope; thus, for a polytrope with index 3.5 the parameter of the nonadiabatic oscillation of the zone would have been 3.8 times as

D. A. Frank-Kamenetskiy

1. Stars--Oscillations 2. Perturbation theory--Applications 3. Oscillations

Card 2/2

ZHEVAKIN, S.A.; TROITSKIT, V.S.; TREYTLIN, N.M.

Atmospheric radio emission and investigation of absorption of centimeter radio waves. Izv.vys.ucheb.sav.; radiofis. 1 no.21 19-26 '58. (MIRA 11:11)

1. Issledovatel'skiy radiofizioheskiy institut pri Gor'kovskom universitete. (Microwaves) (Atmosphere)

20-119-5-18/59 Zhevakin, S. A. AUTHOR: A Uniform Interpretation of Various Types of TITLE: the Pulsation-Dependent Stellar Variability (O yedinoy interpretatsii razlichnykh tipov pul' satsionnoy zvezdnoy peremennosti) Doklady Akademii Nauk SSSR, 1958, Vol. 119, Nr 5, PERIODICAL: pp. 907-910 (USSR) The author shows that on the basis of the calculation of ABSTRACT: the non-adiabatic oscillations of a multilayer discreet spherical model of a stellar shell (in linear approximation) that all known types of pulsation-dependent stellar variability (which are distinguished by phase relations between the oscillations of brightness and the oscillations of the stellar radius) can be obtained from this model with various values of one of the parameters of the model. This parameter is the parameter yzone of the non-adiabaticity of the oscillations of the zone of critical ionization of He II. In the case of the long-period variable stars of the

A Uniform Interpretation of Various Types of the Pulsation-Dependent Stellar Variability 20-119-5-18/59

type O Ceti the beginning of the period of maximum brightness passes the period of the maximum compression of the stars by 1/4 of the oscillation period. Besides there are long-period variable stars in the case of which these two periods approximately coinciding. The last type is called RR Herculis-type according to one of its representatives. The same phase displacement as RR-Her is also displayed by the variable stars of the type B Can Maj. The shell of such stars consists of a zone with double critical ionization of helium and of the stellar atmosphere situated above it. In the case of a given dependence of the non-transparency coefficient of the stellar shell on the density and the temperature (dependent on the chemical composition), the character of the non-adiabatic oscillations mainly depends on the value of the parameter yzone as well as on the value of the ratio matmosphere/Mzone, while the residual parameters are of only little influence. A diagram shows the displacement of the phase Ψ between

Card 2/3

APPROVED FOR RELEASE: 03/15/2001 CIA-RDP86-00513R002064720012-7"

the epochs of maximum brightness and the minimum of the

A Uniform Interpretation of Various Types of the Pulsation-Dependent Stellar Variability 20-119-5-18/59

stellar radius, as well as the ratio between the amplitude of the radiation current on its emission from the star and the maximum of the amplitude of the radiation current which is reached before its entering the ionized zone. After this the various types of stellar reaction are discussed consecutively. There are 1 figure, 1 table, and 22 references, 6 of which are Soviet.

ASSOCIATION:

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PRESENTED:

December 9, 1957, by M. A. Leontovich, Member, Academy of Sciences, USSR

SUBMITTED:

Danamban 4, 1957

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3(1) SOV/20-123-2-11/50 AUTHOR: Zhevakin, 8. A. TITLE: On the So-Called "Relation Period-Luminosity" of Cepheids (O tak nazyvayemom "sootnoshenii period-svetimost' " u tsefeid) Doklady Akademii nauk SSSR, 1958, Vol 123, Nr 2, pp 252-255 PERIODICAL: (USSR) A revision of the relation period-luminosity for the classical ABSTRACT: Cepheids (which was undertaken by the author on a 200-inch telescope because of Baade's (Ref 2) observations) showed the unreliability of statistical daterminations hitherto carried out of the above-mentioned zero from self-motions and from trigonometric parallaxes, and how much optimism had been displayed by the authors of earlier papers when estimating the error limits of the zero values found by them. The author first gives a short report on a theory he developed in 1953. In December 1957 Sandidge (Sandidzh), at the 90th session of the American Astronomical Society, he described his investigations of the period-luminosity ratio of Cepheids. According to Sandidge a spread 1m2 of luminosity corresponds to a given value of the period; thus, instead of a curve period-luminosity, a strip of the breadth 1.2 is obtained. Thus, an unexpectedly good confirm-

经证据的证据,这个是一个是一个是一个证明,但是在这种证明,但是是是是一个证明,但是是是是这种的证据,这一种证明的证明,是一种的证明的证据,这种证明的证据,是是是

On the So-Called "Relation Period-Luminosity" of Cepheids

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ation of the theoretically predicted spread of 1.2 could be given by means of the experimental results obtained by Sandidge. A diagram shows the results obtained by Sandidge for the galactic Cepheids. The author then shows how to explain. Sandidge's results by means of the theory of the variability of stars. His investigations are based on the following assumptions: a) The variable stars have homologous internal structures. b) The shells of variable stars contain one and the same quantity of helium. c) The degree of ionization of He II does not depend on gas pressure and is determined solely by temperature. d) The shells of variable stars are constructed by means of the polytropic index n. In the case of a given period $P \sim \sqrt{R^2/M}$, the relation $R \sim \sqrt{M}$ then holds, where R and M denote the radius and the mass of the stars respectively. For the temperature in the shell it holds that T ~ Mu/R, where u denotes the Emden function with the index n. In the zone of critical ionization it holds that u ~ R/M. Next, it is assumed that the luminosity Lo of the star does not depend on mass, and that the masses of the Cepheids (at last within their various

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groups (short-period-long-period-stc.)) are approximately equal. It then holds that $u_{zone} \sim L_o$. Here u_{zone} denotes the parameter of the non-adiabaticity of the oscillations of the zone of critical ionization of He II in a variable star. With a given period P the following variability of stars of the Cepheid type is possible: $\Delta m = 2.5 \Delta lgL_o = 2.5.0.5 = 1.25$, which agrees well with the value $1.2 cm^2 2.5.0.5 = 1.25$, which agrees well with the value $1.2 cm^2 2.5.0.5 = 1.25$, which agrees well with the value $1.2 cm^2 2.5.0.5 = 1.25$, which agrees well with the value $1.2 cm^2 2.5.0.5 = 1.25$, which agrees well with the value $1.2 cm^2 2.5.0.5 = 1.25$, which agrees well with the value $1.2 cm^2 2.5.0.5 = 1.25$, which agrees well with a given period, Cepheids with great luminosity must have a smaller amplitude of self-oscillations than Cepheids of the same period with a lower degree of luminosity. There are 2 figures and 20 references, 10 of which are Soviet.

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On the So-Called "Relation Period-Luminosity" of Cepheids.

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PRESENTED:

July 2, 1958, by M. A. Leontovich, Academician

SUBMITTED:

June 30, 1958

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SOV/109-4-1-4/30

AUTHORS: Zhevakin, S. A. and Troitskiy, V. S.

TITIE: Absorption of Centimeter Waves in a Stratified Atmosphere (Pogloshcheniye santimetrovykh voln v sloistoy atmosfere)

PERIODICAL: Radiotekhnika i elektronika, 1959, Vol 4, Nr 1, pp 21-27 (USSR)

ABSTRACT: The article gives complete formulae for the evaluation of the absorption in the atmosphere, which take into account the curvature of the Earth and the refraction. The formulae permit the calculation of the absorption for centimetre waves by using the temperature and the absolute humidity at the Earth's surface as the basic data. A path L between points 1 and 2 at heights h₁ and h₂ in the atmosphere is considered (see the figure on p 22), The complete expression for the reflection coefficient at a height h is given by:

$$\mu = \mu_1 \varphi_1(h) + \mu_2 \varphi_2(h)$$
(1)

where x₁ is the absorption coefficient due to the presence Card 1/5

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Absorption of Centimeter Waves in a Stratified . Atmosphere of hydrogen and w2 is the absorption cefficient due to water vapours; ϕ_1^2 and ϕ_2 are functions of height such that $\varphi_1(0) = \varphi_2(0) = 1$ and $\varphi_1(\infty) = \varphi_2(\infty) = 0$. absorption in an element of length dl is equal to e-k(h)d1. $e^{-\kappa(h)dL}$, so that the total absorption is expressed by Eq.(2), where J and J_o are field intensities at any point in the presence of absorption and in the absence of absorption, respectively. Since dL can be expressed by Eq. (5), where n, is the refraction coefficient at point 1 and n is the refraction coefficient at a given point of the path L, the total attenuation coefficient for the wave can be written in the form of Eq.(6). This can also be written as Eq.(7), where quantities l_1 and l_2 are expressed by Eqs.(8). The quantities l_1 and l_2 denote the effective path lengths in a dry atmosphere and in the presence of water vapours. The pressure and the temperature for the standard atmosphere at a height h can be represented by Eqs.(9). On the other hand, the absolute humidity at a height h is given by Eq.(11) Card 2/5 where ρ_0 is the humidity at h = 0 and H₀ is a

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Absorption of Centimeter Waves in a Stratified Atmosphere

characteristic quantity for the water vapour which, for the USSR, is equal to about 2.6 km. The absorption coefficient due to the hydrogen is expressed by Eq.(13) where D is a constant, V is frequency. T is the absolute temperature of the hydrogen, N is the molecular concentration of the hydrogen. AN is the width of the absorption line and $\delta = \Delta N/c$. For the waves of 1.5 to 10 cm the hydrogen absorption coefficient can be simply expressed by Eq.(14). For the standard atmosphere the function ϕ_1 can therefore be expressed by Eq.(18). On the other hand, the function ϕ_2 , which takes into account the absorption due to the presence of water vapours, is expressed by Eq.(22). The final expression for the absorption coefficient is therefore given by:

$$\kappa(h) = \kappa_1 e^{-\frac{h}{H_1}} + a\rho_0 e^{-\frac{h}{H_2}}$$
, (24)

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Absorption of Centimeter Waves in a Stratified Atmosphere where $H_1 = 5.3$ km and $H_2 = 2.1$ km. The quantities L_1 and L_2 can therefore be expressed by Eqs.(27), provided the simplification defined by Eq.(25) is adopted; a_e in Eqs.(27) denotes the effective Earth radius. The quantity L of Eqs.(27) can be expressed by Eq.(29) or Eq.(30); this can also be written as Eq.(31). Similarly, L_2 can be expressed by Eq.(32). Function f(t) in Eqs.(31) and (32) is evaluated in Table 1. Eqs.(31) and (32) were used to determine L_1 and L_2 for various L, O_1 and O_2 . The values obtained are shown in Table 2; these were calculated for $O_1 = 5.3$ km, $O_2 = 2.2$ km and $O_3 = 9000$ km. From the table it is seen that the maximum effective path length in a dry atmosphere does not exceed 274 km and that in the water vapour of the atmosphere is less than 176 km. The authors express their gratitude to

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Absorption of Centimeter Waves in a Stretified Atmosphere

N. M. Tseytlin for his help in this work. The paper contains 1 figure, 2 tables and 6 references; 1 reference is English, 1 French and 4 are Soviet.

ASSOCIATION: Radiofizicheskiy institut pri gosudarstvennom universitete im. N. I. Lobachevskogo, g. Gor'kiy (Radiophysics Institute of the State University imeni N. I. Lobachevskiy, in the town of Gor'kiy)

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SUBMITTED: May 13, 1957.

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